# Life Cycles of Stars



# The universe started with the ...

#### **Big Bang!**

 Everything continued to expand, clouds of dust started to gravitate towards each other forming

stars.



# A star's life begins in a ...

#### Nebula!

 A cloud of gas and dust, consisting mostly of Hydrogen





 Gas and dust begin to clump together to form a Protostar (a baby star).



- The smaller a star is the longer it will live.
  - Larger stars have more fuel, but they have to burn (fuse) it faster in order to maintain equilibrium.
  - Because fusion occurs at a faster rate in massive stars, large stars use all their fuel in a shorter length of time.



 So...A smaller star has less fuel, but its rate of fusion is not as fast. Therefore, smaller stars live longer than larger stars because their rate of fuel consumption is not as rapid.

 The star's main goal in life is to achieve stability, or <u>equilibrium</u>, where pressure from fusion within the core is equal to the force of gravity pushing down on it (this keeps the star "alive").



Continuous steps occur inside the core of a main sequence star, until there is no more Hydrogen.

- Step 1 Nuclear fusion (hydrogen turning to helium).
  Gravity = gas pressure (equilibrium)
- Step 2 Out of fuel
- Step 3 Fusion stops, temperature drops
- Step 4 Core contracts (gravity pulling atoms in)
- Step 5 Increased temperature (more atoms, more collisions) and density in the core reinitiates nuclear fusion, equilibrium is achieved, and the cycle begins again at Step 1.

#### Life Cycle of a Star like our Sun...

Nebula → Protostar → Main Sequence Star → Red Giant → Planetary Nebula → White Dwarf

Protostar	Fusion ignition - Main Squence	Red Giant/Supergiant	White Dwarf/Black Hole
			•
Fetus	Infancy through Adulthood	Middle Age	Old Age-Death

## Life Cycle of a Star like our Sun...

- Our sun is at the Main Sequence stage in its life.
  - When the hydrogen in the core has been used up, the core shrinks and <u>hydrogen fusion</u> begins in the outer layers,
  - which then expands the entire star, turning it into a <u>Red Giant</u>.
  - The sun begins to die when helium is fusing into other elements, then the gases at the sun's surface start to blow away in bursts, called a <u>Planetary Nebula</u> (or halo of gases,
  - Resulting in a hot carbon-oxygen core called a <u>White</u> <u>Dwarf.</u>

#### Life Cycle of a Star With Greater Mass Than Our Sun...

Nebula → Protostar → Main Sequence Star → Red Supergiant → Supernova

Black Hole or Neutron Star



#### Life Cycle of a Star With Greater Mass Than Our Sun...

- Massive stars go through the same life stages as our sun (just on a larger scale) upto the Main Sequence stage,
- Then the massive stars expand into a Red <u>Supergiant</u>,
- Explode into <u>a Supernova</u>,
- Then turn into a <u>Black Hole</u> or a <u>Neutron Star</u>.

#### Life Cycle of Stars

Stars with less than 8 solar masses

8 to 15 solar masses

Stars with 8 solar masses or more

More than 15 solar masses

Modified from Julian Baum, Wildlife Art Ltd

#### Hertzsprung-Russell (HR) Diagram



# **HR Diagram**

- The Hertzsprung-Russell (HR) Diagram is a tool that shows relationships and differences between stars (temperatures, brightness, colors, etc.)
- It is something of a "family portrait." It shows stars of different ages and in different stages, all at the same time.
  - A star in the upper left corner of the diagram would be hot and bright.
  - A star in the upper right corner of the diagram would be cool and bright.
  - The Sun rests approximately in the middle of the diagram, and it is the star which we use for comparison.
  - A star in the lower left corner of the diagram would be hot and dim.
  - A star in the lower right corner of the diagram would be cold and dim.



• The larger a star the brighter the star.

Protostar	Fusion ignition - Main Squence	Red Giant/Supergiant	White Dwarf/Black Hole
			•
Fetus	Infancy through Adulthood	Middle Age	Old Age-Death

http://aspire.cosmic-ray.org/labs/star\_life/hr\_diagram.htm

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